

Monsoons: global energetics and local physics as drivers of past, present and future monsoons

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Monsoons: global energetics and local physics

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Global constraints on momentum and energy govern the structure of the zonal mean trop-26 ical circulation and rainfall. The continental-scale monsoon systems are also facets of a 27 momentum- and energy-constrained global circulation, but their modern and paleo vari-28 ability deviates substantially from that of the longitudinal mean through mechanisms neither 29 fully understood nor well simulated. A framework grounded in global constraints yet encom-30 passing the complexities of monsoon dynamics is needed to identify the causes of mismatch 31 between theory, models, and observations and, ultimately, improve regional climate projec-32 tion. In a first step towards this goal, disparate regional processes must be distilled into gross 33 measures of energy flow in and out of continents and from the surface to the tropopause, so 34 that monsoon dynamics may be coherently diagnosed across modern and paleo observations 35 and across idealized and comprehensive simulations. Accounting for zonal asymmetries in 36 the circulation, land/ocean differences in surface fluxes, and the character of convective sys-37 tems, such a monsoon framework would integrate our understanding at all relevant scales: 38 from the fine details of how moisture and energy are lifted in the updrafts of thunderclouds, 39 up to the global circulations. 40

Most tropical precipitation, whether steady rain or intense showers, falls from cloud clus-41 ters where individual, small-scale updrafts are organized over a few hundred kilometers in dis-42 crete weather systems. These, in turn, are orchestrated by planetary-scale circulation features: the 43 monsoons and the Inter-Tropical Convergence Zone (ITCZ, Figure 1a and b). The clustering of 44 individual clouds is the visible signature of an otherwise invisible global stirring. The notion that 45 a link exists between smaller and larger scales underpins our understanding of tropical rain belt 46 dynamics and is the basis of their representation in global climate models. Our understanding, 47 however, remains incomplete, as evidenced by our inability to achieve reliable predictions of how 48 the ITCZ and monsoons respond to external forcings. 49

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Without the benefit of outdoor controlled experiments, or the possibility of validating pre-

dictions of long-term future changes, the reliability of our prediction tools must be tested against past records. Confidence in projected climate responses increases when dynamic theories built on contemporary observations also explain past conditions and when simulations skillfully reproduce the paleo record.

Good tests of our theories and models are the Holocene waxing and waning of monsoons 55 in response to orbitally-driven changes in incoming solar radiation^{1,2} (Figure 2). During the early 56 to mid Holocene (11,000-5,000 years before present), the Northern Hemisphere received more 57 insolation during its summer than today, while the Southern Hemisphere received less during its 58 summer. In the modern climate, monsoons export energy from regions where the sun delivers 59 most-so we would expect mid-Holocene monsoons to be stronger in the Northern Hemisphere 60 and weaker in the Southern Hemisphere, compared to today. Paleoenvironmental data from that 61 period indeed indicates increased rainfall in North and Central America^{3,4}, a stronger Indian mon-62 soon and increased inland penetration of the monsoon into China^{5,6}, and a spectacular rainfall 63 expansion in northern Africa^{7,8}: abundant lake, pollen and archaeological evidence documents 64 wetting and vegetation increase over much of the Sahara. 65

While qualitatively consistent with our expectations, the major expansion of the Northern 66 Hemisphere monsoons is widely underestimated in climate model simulations^{1,2}, in particular 67 over the Sahara⁹ (Figure 2b). This bias can be reduced by modeling earth-system feedbacks¹⁰ 68 or by imposing the observed changes in vegetation, surface water storage, wetlands, soils, and 69 mineral dust as boundary conditions¹¹⁻¹³. Yet, simulations that organically produce both the rain-70 fall distribution and the vegetation types that are consistent with the records still elude us². Moist 71 atmospheric dynamics and its coupling with other aspects of the Earth system, including vegetation 72 cover and soil properties, remain prime suspects for the failures. Moreover, the sign of observed 73 Southern-Hemisphere changes is not fully consistent with the predictions of reduced rainfall from 74

⁷⁵ both arguments based on the hemispheric summertime insolation forcing and complex model sim-⁷⁶ ulations. Although many more quantitative estimates of rainfall anomalies from the Southern ⁷⁷ Hemisphere are needed to paint a full picture, many palaeoenvironmental records from South-⁷⁸ ern Africa imply *increased* precipitation¹⁴ and palaeoenvironmental records from South America ⁷⁹ and Australia seem to show (their interpretation being somewhat controversial) mixed wetting and ⁸⁰ drying signals^{15, 16}. Thus, neither complex models nor theoretical intuition are sufficient to explain ⁸¹ past monsoon records.

In a traditional dry paradigm, monsoon circulations are akin to continental-scale land-sea breezes driven by surface *temperature* contrast, their strength increasing with the contrast. But this is too simplistic. Monsoon lands are hottest before the start of the monsoon, but the circulation is strongest in late summer (when increased rain and cloudiness have cooled the land and reduced the contrast). In future climate projections, land-sea temperature contrasts universally strengthen, but monsoon circulations generally weaken¹⁷ as does, in some instances, regional early-season rainfall^{18,19}.

A better paradigm views monsoons not as giant heat lows for which rainfall is a side effect, or as circulations "driven by" the latent heating of rainfall, but as moist energetically-direct circulations tightly coupled to precipitating convection; a facet of the general overturning of the tropical atmosphere inextricably linked to the Hadley circulation and the zonal mean ITCZ²⁰.

At seasonal time scales, convection acts to release any column instability and to bring the free-tropospheric temperature in line with the moist static energy (MSE) of the boundary layer below it²¹. Horizontal atmospheric motions homogenize the free-tropospheric tropical temperatures to one vertical profile which reflects the conditions of the major convective centers. In this view, peak rainfall should coincide with peak low-level MSE; cooler and dryer surfaces can support only shallower or suppressed convection. Land and ocean observations confirm this theoretical

prediction²². Meanwhile, the upper-level divergence over these convective regions (maxima of 99 low-level MSE) implies that the circulations associated with peak rainfall export total energy away 100 from the centers of deep convection. In other words, monsoons and the ITCZ are components of 101 a planetary energetically-direct circulation that links land and ocean rainfall (similar to a "global 102 monsoon"²³) and shows coherent variability from seasonal^{24,25} to geological time scales^{20,26,27}. 103 Such an encompassing view is an important theoretical advance. Its formulation for the zonal 104 mean tropical circulation-known in the literature as the energy-budget framework-underpins our 105 understanding of why the ITCZ moves meridionally in response to forcings that originate well 106 outside the tropics^{20, 28, 29} (Section 1). 107

Further progress in understanding and simulating regional rainfall anomalies requires us to 108 extend the energy-budget framework developed for the Hadley circulation to describe how the full 109 tropical circulation redistributes energy vertically and horizontally, from its poleward boundaries 110 and across the tropics. The extension of the energy-budget framework from a purely zonal for-111 mulation to one that includes zonal asymmetries would parallel the progress that has been made 112 in understanding how the momentum budget constrains tropical circulations. The momentum-113 budget framework has led us from a starting view of the Hadley circulation³⁰ and-to a degree-114 the monsoons³¹ as axisymmetric angular-momentum-conserving circulations, to the recognition 115 that momentum transport by eddies is crucial to maintaining the zonal mean circulation in most 116 cases^{32,33}, to a deepening understanding of the role of the vorticity transport by stationary ed-117 dies in both localizing the monsoon circulation and regulating its intensity^{34,35}. An analogous 118 energy-budget framework would invoke zonal asymmetries in surface properties, which introduce 119 horizontal gradients in the distribution of energy, amplify the importance of both stationary and 120 transient eddy transports, and can generate shallow circulations which can be energetically in-121 direct (Section 2). These asymmetries are expressed in the observable differences in convective 122 weather characteristics between land and ocean (Figure 1c and d), which aggregate into different 123

ascent profiles and vertical energy transport (Section 3).

The eventual goal is to combine the momentum-based and energy-based theories of tropical 125 circulations and rainfall into a self-contained model for the tropical climate, one that describes how 126 the interplay between energy and momentum fluxes in a moist atmosphere-where a profusion of 127 weather phenomena, cloud types, and scales of motion is organized in mean and eddy effects— 128 drives the seasonal evolution of oceanic and continental rainfall and controls monsoon diversity, 129 variability, and response to external forcing. Understanding how the clouds within such circula-130 tions modify the energy input into the atmosphere through changes in both radiative fluxes and 131 turbulent surface fluxes would be the next challenge $^{36-38}$. In the rest of this paper, we propose a 132 first step towards our goal: an energy-budget framework suitable for monsoons. 133

134 **1** The explanatory power of energetic constraints.

Energetic constraints provide a parsimonious explanation of how the zonal-mean tropical rainfall 135 shifts its position in response to internal variability and external forcings^{20,28,29}. The ascending 136 branches of the Hadley cells are in the deep tropics, where the solar radiation absorbed by the 137 earth most greatly exceeds the terrestrial radiation emitted to space. Energy is transferred from 138 the surface to the atmosphere by fluxes of radiation, sensible heat and moisture (latent heat) and a 139 planetary circulation moves this excess energy towards high latitudes (Fig. 3a). Moist surface air 140 converges and rises in the ITCZ, thus cooling adiabatically, condensing moisture, and forming pre-141 cipitating clouds. As a result, maximum rainfall is broadly co-located with the boundary between 142 the Hadley cells (Fig. 3c). 143

These ideas can be formalized by focusing on the atmospheric energy budget in the annual and zonal means^{20, 28}. Averaging in time and longitude disposes of tendency terms and zonal fluxes, so that, under the assumption that eddy terms are unimportant, net energy input into the atmosphere is balanced by the divergence of the vertically integrated energy flux by the Hadley circulation \mathcal{F} and the ITCZ coincides with the "energy-flux equator," i.e., the latitude at which the energy flux $\mathcal{F} = 0$ and changes direction.

The Coriolis force is weak in the tropics and atmospheric waves are effective in smooth-150 ing out pressure gradients. Because the tropical troposphere cannot maintain strong temperature 151 gradients, an extra-tropical eddy heat flux that reaches the tropics will be carried into the tropics 152 and through-to the mid latitudes of the other hemisphere. This argument appears to apply more 153 broadly: any asymmetry in the energy flux across the northern and southern edges of the trop-154 ics is felt throughout the tropical band³⁹, making the position of the energy-flux equator sensitive 155 to extra-tropical forcings²⁰. The same argument explains how the inter-hemispheric asymmetry 156 in seasonal insolation drives the north-south annual migration of the ITCZ²⁴. However, because 157 the cross-equatorial Hadley cell is always the strongest throughout the annual cycle, and because 158 maximum ascent and rainfall are concentrated within the cross-equatorial cell and equatorward of 159 the energy-flux equator, seasonal rainfall shifts are less pronounced than those of the energy-flux 160 equator²⁴ (Fig. 3a). In comprehensive models, changes in the annual-mean position of the ITCZ, 161 their inter-model spread, and the seasonal migration of the rain belt have been shown to follow 162 the same quantitative relationship with energy transports, albeit with some scatter and uncertainty 163 (a 1PW change in the cross-equatorial energy flux leads to about a 3° shift in the position of the 164 ITCZ, Fig 3b). This correspondence has been used to suggest that the same dynamics control both 165 the meridional shifts characteristic of the annual cycle and the variability of mean rainfall at paleo 166 time scales^{20,24,40}. 167

This theory for the ITCZ position sees the tropical rain belt as the expression of the conservation law that governs energy in the climate system; it links the occurrence of convection not just to the local environment, but to the planetary adjustments that bring the global atmosphere towards

equilibrium; and it subsumes arguments that link the ITCZ position to gradients in tropical SST. It 171 has been used to explain the location of the modern ITCZ north of the equator as a consequence 172 of energy transport by the thermohaline ocean circulation⁴¹, the southward shift of tropical rainfall 173 during the last glacial maximum and Dansgaard-Oeschger events as a consequence of northern 174 high-latitude cooling^{29,40}, the effect of Eurasian afforestation on monsoons in the mid Holocene⁴², 175 the Sahel drought in the 1970's and 1980's as part of a global-scale southward ITCZ shift due 176 to sulfate aerosols⁴³, and the role of Southern Ocean heat uptake in setting up the hemispheric 177 asymmetry in future tropical rainfall changes⁴⁴. 178

However, the existing framework is limited in important ways. First, it is incomplete as 179 a predictive theory because internal radiative feedbacks from clouds and water vapour can over-180 whelm the external forcings in setting energy gradients, are not easily predicted from gross en-181 ergetic constraints, and vary substantially between models^{36, 37, 45}. Even when the radiative effect 182 of clouds is carefully controlled in models^{46,47}, changes in oceanic heat transport can oppose the 183 inter-hemispheric difference in energy input in the atmosphere and complicate the response of 184 the ITCZ⁴⁸ to external forcings. Indeed, the tight coupling between atmospheric and oceanic heat 185 transport^{20,49–51} suggests an expansion of atmospheric-only arguments. Second, changes in tropical 186 rainfall are often better described as intensifications⁵² or contractions ^{19,53–56} of the climatological 187 net rainfall pattern, rather than shifts, and some regional anomalies are strongly driven by localized 188 gradients in surface conditions ^{57–59}, including between ocean and land^{60,61}. Third, the assumption 189 that rainfall is the product of deep convection in which ascent extends throughout the troposphere 190 and maximizes at mid levels is crucial to the portrayal of the Hadley cell as energetically direct. 191 Quantitatively, this requires that the same large-scale circulation that converges moisture into the 192 ITCZ also diverges enough static energy from the ITCZ at upper levels that the vertically-integrated 193 result is a net export of MSE—a result that depends heavily on where exactly in the column air is 194 converging and diverging. Yet, rainfall is also produced in circulations with shallow components 195

and bottom-heavy ascent profiles (see also Fig 4), and direct calculation of the energy flux out of 196 such regions indicate a net import of energy^{62,63}. Fourth, the assumption in the ITCZ energy-budget 197 framework that eddies are not important is debatable. Even well within tropical latitudes, transient 198 eddies transport latent heat poleward⁶⁴ and thus might be as important as the zonal circulation in 199 transporting energy, especially where the latter is weak. While the bulk effect of transient eddies 200 could be subsumed into the framework developed for the ITCZ in an aquaplanet^{28,65}, the presence 201 of land introduces stationary eddies^{34,66} and localizes transient eddies³¹ in ways that preclude the 202 straightforward application of the zonal mean framework to limited longitude bands (Fig 2c). This 203 is particularly true in a moist atmosphere, where feedbacks between surface fluxes, clouds, and 204 latent and radiative heating can amplify and extend the asymmetries in the forcing⁴⁵. These limi-205 tations hamper our ability to use the existing energy-budget framework for the ITCZ to accurately 206 predict the response of regional rainfall to past or future forcings. Yet the fundamental insight that 207 regional changes are expressions of global conservation laws should not be abandoned. 208

209 2 The need for an energy-budget framework for monsoon systems.

Zonal asymmetries and contrasts between land and ocean are fundamental to the energy and the 210 momentum budget of monsoon circulations^{13,34,67,68}. To close, the energy budget must include 211 the effect of zonal transports and of the complex vertical structure of the meridional circulation, 212 both of which are the result of inhomogeneities in surface properties. Inhomogeneities do oc-213 cur over oceans (such as between warm pools and the equatorial cold tongues), but gradients in 214 surface properties are especially strong at coastlines and over continents due to orography, vari-215 ability in the characteristics of soils and vegetation, and-in a positive feedback-to the response 216 of the land system to differences in precipitation. The conceptual model of the Hadley cell as a 217 simple deep overturning meridional circulation (Fig. 3c) can broadly capture the main features of 218 the ITCZ, but is insufficient to describe the more complex monsoon circulations. In many regional 219

monsoons^{22,69}, gradients in surface temperature and sensible heat fluxes (e.g., between the hot sub-220 tropical desert and the cooler oceans and equatorial rainforests) force a shallow circulation with 221 dry ascent poleward of the monsoon rainfall. This circulation includes the low-level monsoon flow 222 that fuels moisture to the rainfall band and the dry return flow above that can cap deep convection, 223 affecting the frequency of rainfall and the occurrence of severe convection (Figure 1c,d and 3c). 224 Re-evaporation of rainfall in a dryer lower troposphere (Fig. 3c, cloud types) can also affect the 225 vertical distribution of latent heating and, concurrently, ascent (see also the next section). The 226 presence of land, thus, changes the vertical transports in the meridional divergent circulation. The 227 strong horizontal gradients in temperature and moisture associated with surface type also induce 228 strong non-divergent, rotational flows that transport energy and moisture horizontally (Fig. 3c, 220 broad arrows). Past seminal work⁶⁷ has shown that ventilation, the transport of low-MSE oceanic 230 air by the rotational flow, is key to setting the poleward extent of the monsoon rainfall. Recent 231 studies with more comprehensive models have confirmed the important role of the rotational flow 232 in balancing the energy budget of monsoon regions⁷⁰ and in driving the onset of off-equatorial 233 rainfall³⁴. The annual changes in surface properties that take place during the progression of the 234 rainy season are also reflected in changing flows of energy in the atmosphere. The most obvious 235 change is soil water content, which affects evaporative fluxes and albedo and, because of its high 236 spatial variability, can introduce sharp gradients in surface properties at small spatial scales⁷¹. 237

There is a vast literature that focuses on the individual regional monsoons and emphasizes the zonally asymmetric regional flow and its interaction with mountains and regional oceans. There is value in formulating such regional analysis within a MSE framework. For example, the extent and intensity of the South Asian monsoon depends on the presence of the Himalayan mountains because the latter shield the MSE maximum over India from low-MSE extratropical air⁷². In another example, the African monsoon appears to be particularly susceptible to moisture anomalies entering from the North Atlantic^{73–76} and the Mediterranean⁷⁷ (even though the main moisture source for the monsoonal rains is in the tropical ocean) because they regulate the depth of convection,⁷⁸ which is in turn linked to energetic requirements of the large-scale flow⁷⁰. How the profile of convective ascent evolves across the monsoon season modulates such sensitivity, by making changes in either the boundary layer MSE (such as the horizontal advection or recycling of moisture) or the free-tropospheric dynamics (such as changes in upper-level vertical stability) more or less relevant (a similar dependence of rainfall sensitivity on low-level and upper-level processes is also seen across models^{70,73}).

The broad-brush framework built around the MSE budget has been successful in explain-252 ing why and how processes as disparate as aerosol microphysics and the oceanic thermohaline 253 circulation can affect the zonal mean tropical rainfall. Here, we advocate extending the energy-254 budget framework from its current form, appropriate to the ITCZ/Hadley cell, to one appropriate 255 for monsoons: one that quantifies how the presence of zonally-confined land masses modifies 256 both the input and the flow of energy in the atmosphere and thus changes the leading terms of 257 the vertically-integrated atmospheric energy budget. In this section we have emphasized (i) how 258 land/ocean differences in surface properties change the input of moist and dry energy to the atmo-259 sphere, (ii) how the continental-scale stationary circulations that develop in response advect MSE 260 gradients horizontally through the geostrophic flow, and vertically through shallow circulations, 261 and (iii) how changes in the dominant balance shape both the diversity of regional monsoons and 262 monsoon evolution through the development and decay of the rainy season. In the next section we 263 focus on the role of vertical motion at the scale of clouds, and how it shapes the vertical fluxes of 264 moist energy and its input in the atmosphere. 265

3 The diversity of convection in tropical precipitation systems: interaction with the large scale climate.

Recent observations have highlighted the rich diversity of convective systems (Figs. 1 and 4), and 268 have allowed fundamental insights into the processes that govern them. Convective precipitation 269 appears controlled by both low-level and deep-column moisture, which set the buoyancy of en-270 training ascending parcels. Thus, while early theories predict the position of the rainfall maximum 271 from just the boundary-layer MSE (Sec. 1), full tropospheric water vapour is key to fully account 272 for the spatial and temporal variability of rainfall intensity. In turn, processes such as detrainment 273 from precipitating clouds and re-evaporation of rain make the tropospheric water vapour depend 274 on the occurrence of rainfall. This two-way coupling underpins the correspondence between rain-275 fall amounts and total humidity found over land and ocean (see also Fig 1a) and is encapsulated 276 in an exponential relationship of daily precipitation intensity on the integrated humidity⁷⁹ (with 277 some variations across convective system⁸⁰ and their drivers⁸¹). However, rainfall characteristics, 278 such as intensity, organization, and duration, and the vertical and temporal distribution of clouds 279 depend on factors other than column humidity, including wind shear, the larger-scale flow, and the 280 properties of the surface boundary^{71,82,83}. Continental updrafts are often deeper and more intense 28 than oceanic updrafts (Fig. 4a,b), as evidenced by the preferential occurrence of lightning over 282 land (Fig. 1c), but land convection varies greatly through the day (cf. Figs. 4 c and d) and the 283 season in depth, organization, and lifetime—affected by surface inhomogeneity and by stronger 284 triggering and inhibition processes. 285

The growing appreciation of the diversity of convective cloud systems has yet to mature into enough understanding of the interplay between clouds and large-scale dynamics to create muchneeded convective parameterizations able to describe such diversity. Parameterizations are still overly reliant on so-called quasi-equilibrium formulations: the gross effects of convection are taken

to be nearly instantaneous and deterministic functions of large-scale forcings, thermodynamics 290 drivers are overemphasized over dynamics, and convective organization is typically ignored^{84,85}. 291 The approximations lead to a preference for tall, disorganized convection and the lack of both 292 shallower, developing convection⁸⁶ and more persistent cloud systems and is likely the source of 293 climate models' systematic errors in rainfall timing (for example, the too-early peak in the diurnal 294 cycle of land rainfall). It affects the climate at longer time scale as well: model biases in seasonal 295 rainfall are typically established within few days after initialization from observations, and only 296 later amplified in the coupled system, pointing to the dominance of fast atmospheric processes⁸⁷ 297 in setting the bias. Such long-ranging effects are not surprising: the aggregate effect of convection 298 is reflected in the profiles of horizontal convergence and divergence and in cloud and moisture 299 radiative effects and will thus affect both the net vertically-integrated energy flux and the total 300 energy input into the atmosphere. The extent and vertical distribution of clouds modify the net 301 energy in the atmosphere via changes in radiative fluxes and, by modifying the structure of the 302 boundary layer, turbulent fluxes. The vertical profile of ascent modulates the energy transport by 303 the circulation, to the point that a predominance of bottom-heavy or top-heavy convection can 304 determine whether the circulation imports or exports MSE to or from a convective region^{62,63}. 305 Thus, shortcomings in the mix of deep convective and stratiform rain production, of warm and 306 cold cloud microphysics, and of mixing and re-evaporation, can translate into global scale biases⁸⁸ 307 and introduce another source of uncertainty in future projections⁷⁰. 308

High-resolution cloud-resolving dynamical models can portray the full evolution and organization of cloud systems^{88,89} and—when coupled to parameterized large-scale fields⁹⁰ that describe a broad suite of boundary conditions and environments— are being used to investigate how the environment shapes the rich diversity of cloud characteristics and how cloud processes feed back on the environment. These experiments have shown that the specific way in which land convection evolves during the day (such as through morning fog or land-sea breezes) is a key determinant of the large-scale mean and the seasonal evolution of rainfall and other environmental variables^{91,92}. More research on tropical land convection is needed to elucidate whether this cross-scale link is achieved because land processes and clouds modify the energy input into the atmosphere, or because the daily evolution of convection changes the profile of ascent and this effect is rectified into changes in the vertical energy flux.

320 **4** Synthesis

Common biases in the simulation of the diurnal and seasonal cycle of rainfall highlight structural 321 deficiencies across global climate models. Moreover, common biases in the simulation of rainfall 322 in past climate states indicate that common structural deficiencies also affect the modeled responses 323 to changes in forcing. This diminishes the value of consensus in projections of climate response to 324 anthropogenic forcings as an indication of reality. Convective parameterizations, which produce 325 most tropical rainfall within current climate models, have inherent limitations that are a likely cause 326 of bias, but whether cloud-resolving global models will be able to adequately capture observed 327 rainfall variations in modern or past climate states remains to be seen. Oceanic processes that 328 amplify atmospheric biases in the seasonal cycle⁹³ are likely to play a role in setting long-term 329 trends⁹⁴ and paleo simulations highlight the additional importance of earth-system feedbacks, such 330 as between precipitating atmospheric dynamics, vegetation cover, and soil composition. Having a 331 small set of metrics by which to characterize simulations of a climate phenomenon would make 332 the task of model development less daunting; a modern energy-budget framework, one modeled on 333 past efforts^{20,21,32,95,96} but aware of the advances and challenges outlined in this review, can supply 334 such metrics for tropical rainfall. 335

³³⁶ Developing the framework we are advocating will require the combination of observational, ³³⁷ modeling and theoretical approaches. Both modern and paleo observations must be the ultimate

tests of our theories. Fine-resolution, high-frequency observations of rainfall and of the vertical 338 atmospheric structure, together with cloud-resolving simulations, have the potential to shed much-339 needed light on how convection becomes organized and how monsoon rainfall interacts with its 340 environment from sub-daily to monthly timescales⁸⁴. To be truly useful for benchmarking met-341 rics and developing parameterizations, current efforts must be expanded to encompass a broader 342 cross section of tropical environments and climate regimes—including a variety of chronically 343 undersampled land regions. Collection of new paleo evidence and quantitative reconstructions of 344 precipitation from the Southern Hemisphere will refine our picture of paleo monsoons in the mid 345 Holocene and provide a well defined target for testing model simulations and their interpretation 346 within this common framework. The opposite is also true: forward proxy models and climate 347 models are needed to guide the interpretation of data, morph sparse environmental observations 348 into a coherent portrait of past climates, and help prioritize new data acquisition. Multi-model 349 simulations⁹⁷, simulations with a broad set of forcings⁴⁸, and simulations with models of different 350 vintage are needed together to provide a robust test for a new monsoon framework. 351

Drawing our lessons from the development of budget-based theories of the ITCZ, we pro-352 pose that a systematic understanding of the monsoons will require a comprehensive hierarchy of 353 model simulations and a common set of energy and momentum diagnostics to compare results 354 across the model hierarchy and with observations. A key element of such a hierarchy are idealized 355 model setups in which energy is conserved through fluxes at the atmospheric boundaries and that 356 are designed to highlight the effect of different land characteristics—themselves isolated through 357 ad-hoc idealizations—on the response of the tropical rain belt to a range of external forcings^{98–100}. 358 Accurate diagnostics of how energy flows into and within the tropical atmosphere, and via differ-359 ent elements of the atmospheric circulation (from the zonal mean, to stationary eddies, to synoptic 360 eddies, to much faster variability at the scale of convection) are often difficult to calculate from 36 available model output⁵⁴. Yet they must underpin any metrics used to elucidate which small- and 362

large-scale processes are essential to the monsoon systems (and by which interactions); to test the 363 coherence of our dynamical theories; and to benchmark model development. A theoretical model 364 that can deduce the behavior of tropical rainfall from simple indicators based on conservation laws 365 of the physical system with minimal empirical or ad-hoc assumptions would cap the hierarchy and 366 provide the grounds for interpreting both idealized and realistic simulations and for understand-367 ing the observed behavior of the earth system. To derive a unified theory of the ITCZ and the 368 regional monsoons from basic conservation laws is a formidable challenge, but we must meet it to 369 confidently link past to present to future. 370

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Authors' Contributions MB lead the writing process and produced Figure 1 (from an idea by BEM), Figure 2 (from data provided by SPH and PB), and Figure 3 (in collaboration with WRB and AV). CS produced Figure 4. All authors collaboratively drafted the outline of the paper and greatly contributed to the writing process.



Figure 1: (a) Rainfall (color) on July 27th, 2015 and the high atmospheric moisture enveloping it (indicated by the 45mm contour of column-integrated water vapour, the full field is in gray); (b) zonal-mean rainfall for the same day (dark green) and climatological values for the same period (light green). (c) July climatological mean intensity of instantaneous near-surface rainfall (from TRMM precipitation radar, units of reflectivity) and occurrence of lightning (red dots) on July 27th, 2014 (ascending passes of the Lightning Imaging Sensor on TRMM). (d) zonal mean rainfall intensity for land and ocean regions. See on-line method section for further details.



Figure 2: (a) Mid-Holocene mean annual precipitation (MAP) anomalies (color circles, sized by the number of used reconstructions). (b) CMIP5/PMIP3 simulated (gray circles) and reconstructed (black circles with error bar) mid-Holocene MAP changes and grid cells contributing to the reconstruction and model means (numbers at right) for northern Africa latitude bands. (c) Correlation of decadal rainfall variations with the first principal component of the zonal-mean precipitation in transient Holocene simulations. See methods for data references and details.



Figure 3: (a) Zonal-mean rainfall (dashed), ITCZ position (green), and zonally- and vertically-integrated atmospheric energy transport (in PW, shaded). (b) Climatological ITCZ latitude as a function of the vertically-integrated atmospheric energy transport at the equator. Numbers correspond to calendar month. The ellipse is sloped at 3° ITCZ shift per 1 PW energy flux. (c) ITCZ (blue background) and monsoon (green background) schematics. The Hadley cells (dark solid lines) meet in the northern tropics, maximum ascent and rainfall occur in the winter cell (equatorward of the cell boundary) close to maximum low-level MSE (darker surface shading). Monsoons are distinguished by the additional shallow meridional circulation (dashed) and ventilation by the rotational horizontal flows (ribbon arrows) and by a different distribution of cloud types. See on-line methods.



Figure 4: Cloud characteristics (top) and ascent profiles (bottom) over the eastern Atlantic ITCZ (left) and western Africa (right) as function of height. (a-b) The Contoured Frequency with Altitude Diagrams (CFAD, filled contour) shows the frequency (logarithmic scale) of storm-top height as a function of rainfall intensity (units of reflectivity, dBZ). Contours: percentile lines (solid=median, dashed-dotted = 99th percentile) of reflectivity frequency. Frequency is normalized at each level and all rain types are included. (c-d) Daily evolution of the vertical velocity profiles (pressure coordinates; negative values indicate upward motion) from 6-hourly MERRA reanalysis. See on-line method section for details.

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Figure 1: Characteristics of observed tropical rainfall. On any given day, rain is produced by cloud 643 systems that range in size between individual convective towers (1-10km) to mesoscale convective 644 systems and tropical cyclones (> 100 kilometers). Despite substantial variability on short time 645 scale and small spatial scales, most disturbances are organized within the large-scale rain belts 646 formed by the monsoons and the inter-tropical convergence zone (> 1000 km). The map (a, color) 647 and zonal mean (b, dark green) of daily rainfall are obtained from merged GPM satellite mea-648 surements calibrated against rain gauges (Huffman et al., 2015). Column integrated water vapour 649 (a, grey shading and magenta contour) was obtained from the ERA Interim reanalysis product 650 (Dee et al., 2011). The summer climatological zonal mean rainfall (b, light green) is calculated as 651 the 1998-2014 average of the 11-day period centered around July 27th and is obtained from the 652 TRMM-3B42 rainfall estimate (Huffman et al., 2007). Climatological conditional intensity of rain-653 fall (c, color shading, and d) is obtained from the precipitation-radar data on the TRMM satellite 654 (Biasutti et al, 2011) and is expressed as a reflectivity in units of dbZ (decibels of Z). Reflectivity is 655 the amount of transmitted power returned to the radar receiver; light rain is detected by the TRMM 656 PR when the dBZ value reaches 18 (corresponding to about 0.4mm/hr). The higher the dBZ, the 657 stronger the rainrate. Uncertainty in the methods for translating reflectivity into a quantitative 658 precipitation estimates are detailed in Villarini and Krajewski (2009). Lightening flashes for July 659 27th, 2014 are obtained from Lightning Imaging Sensor during ascending passes of the TRMM 660 satellite (https://lightning.nsstc.nasa.gov/lisib/lisbrowse.exe?which=qcyear=2014day=208). 661

<u>Figure 2:</u> Changes in rainfall in the Holocene indicate a complex behavior across the monsoon systems, not fully captured by climate models, and do not suggest that meridional displacements of the zonal mean ITCZ explain a large fraction of the variance of tropical continental rainfall. (a) The expectation that Southern Hemisphere monsoons would be weaker in periods of

weaker SH summer insolation is not fully supported by current observations. Quantitative recon-666 structions of changes in mean annual precipitation (MAP) between the mid Holocene (11,000-667 5,000 years before present) and present day in colored circles, the color indicates the size of the 668 anomaly (mm, colorbar) while the size indicates the number of reconstructions used for the es-669 timate, as indicated by the legend inside the map. The data was first published by Bartlein et al 670 (2011). (b) CMIP5/PMIP3 simulated and reconstructed changes in mean annual precipitation in 671 the mid Holocene for 5° latitude bands across northern Africa (longitude 20°W to 40°E between 672 0 and 45° N), where the model results are averages for the grid cells with observations and each 673 model is represented by a different gray circle. The mean and standard error of the reconstruc-674 tions is shown in black and the number of grid cells contributing to the reconstruction is shown 675 for each latitude band. CMIP5/PMIP3 data were available thanks to the World Climate Research 676 Programme's Working Group on Coupled Modelling, which is responsible for CMIP. For CMIP 677 the U.S. Department of Energy's Program for Climate Model Diagnosis and Intercomparison pro-678 vides coordinating support and led development of software infrastructure in partnership with the 679 Global Organization for Earth System Science Portals. The figure is modified from Perez-Sanz et 680 al., 2014; (c) Correlation maps (shading) of decadal rainfall variations at each gridpoint with the 681 first principal component of the zonal mean precipitation in the fully-forced TrACE-21000 (Otto-682 Bliesner et al, 2014) simulation for the period 9.5ka to 0.5ka before present. The dashed line is a 683 representative contour for the mean precipitation and indicates the climatological position of the 684 rain belt. 685

<u>Figure 3:</u> The energy-budget framework for the tropical rain belt. (a) The seasonal evolution of the zonal mean observed climatological rainfall (dashed contours, only the 4, 6, and 8mm/day isolines are shown) and ITCZ position (defined as the centroid of zonal-mean rainfall within 20° N/S; green line) are superimposed on a reanalysis-based estimate of the zonally and vertically integrated atmospheric energy transport (shaded, warm and cool colors indicate north⁶⁹¹ ward and southward transport, respectively, and the white area indicates the energy flux equator, ⁶⁹² units of PW). (b) The seasonal relationship between the vertically integrated atmospheric energy ⁶⁹³ flux at the equator (\mathcal{F}_o) and the ITCZ position. The energy flux associated with the net mass move-⁶⁹⁴ ment between the hemisphere is retained here, leading to a phase lag between the two fields. The ⁶⁹⁵ slope of the relationship, given by the direction of the major axis of the ellipse, indicates a 3° shift ⁶⁹⁶ for a 1 PW energy flux, consistent with calculations that omit the barotropic circulation.

The climatological rainfall in Figure 3a,b is calculated as the 1979-2013 average of the Climate Prediction Center (CPC) Merged Analysis of Precipitation (CMAP) dataset (Adler et al., 2003). The energy fields are from ERA Interim reanalysis (Dee et al., 2011) over the same period. The cross-equatorial energy transport is calculated eliminating the mass budget residual before vertically integrating the fluxes, the data was provided by the National Center for Atmospheric Research (retrieved from https://climatedataguide.ucar.edu/climate-data/era-interimderived-components).

(c) Schematic of the ITCZ (depicted over an oceanic surface, blue half on the left) and 704 monsoon (depicted over a continental surface, green half on the right) circulations for northern-705 hemisphere summer. The summer and winter Hadley cells (dark solid lines) meet in the northern 706 tropics, close to where the low-level moist static energy (MSE) is maximum (darker shading at 707 the surface), consistent with convective quasi-equilibrium theory. Most upward motion and thus 708 most rainfall occurs in the ascending branch of the stronger, winter cell, so maximum rainfall is 709 slightly equatorward of the Hadley cell boundary and the energy-flux equator. As is the case for 710 the ITCZ, rainfall associated with the monsoonal circulation is positioned slightly equatorward of 711 the maximum surface moist energy (dark blue shading) and is associated with large-scale ascent in 712 local meridional overturning cells whose strength is greater when the ascent is further away from 713 the equator. Key distinctions for the monsoons are in the complexity of the circulation and the 714

distribution of cloud types. Notice in the land portion of the diagram the presence of a shallow 715 meridional circulation (dashed lines) with ascent poleward of the rain band and a dry return flow, 716 the rotational circulations associated with the low-level cyclone (light and dark blue ribbon-width 717 arrows indicate negative and positive transport of MSE), mid- and upper-level land anticyclones, 718 and the oceanic anticyclone (anticyclonic circulations are depicted with ribbon-width grey arrows). 719 Notice also the deeper and more intense convection over land (indicated by a distribution of clouds 720 that include more overshoots and fewer clouds lacking an anvil), more lightning, less rain from 721 warm cloud (fewer clouds without anvil), and more re-evaporation of rain (dotted rain from the 722 anvil cloud). 723

This schematic highlights those aspects of an hypothetical "essential" monsoons that are addressed in this paper. It is not meant to represent any particular monsoon system, as each is highly affected by the geometry of the continent, the location and orientation of orography, the geographical distribution of surface types, including deserts, and oceanic processes unique to each ocean basin.

Figure 4: Cloud characteristics and ascent profiles over ocean and land during the peak of 729 the rainy season. Left: the eastern Atlantic ITCZ (40-20°W 5-12.5°N). Right, western Africa 730 (10°W-10°E 7.5-15°N). The Contoured Frequency with Altitude Diagrams (CFAD) in the top pan-731 els show that western Africa has deeper, more intense convective cells while the eastern Atlantic 732 has more mid-level rainy cloud. The filled contour shows the log of the frequency of storm top 733 height as a function of radar reflectivity (a measure of rainfall intensity). Both regions show that 734 storms reaching 5km in height and measuring reflectivities of less than 30dBZ are the most fre-735 quent, but this peak is more pronounced over ocean than land. Conversely, the land region shows 736 more frequent instances of storms that have intensity above 50dBZ (colors extending to the right 737 of the diagram) and reach 18km in height (colors extending to the top of the diagram). Overlaid 738

on the color are percentile lines of reflectivity frequency at each level with the black line equal 739 to the median at each level. The far right line is the 99th percentile line. Notice that the median 740 surface reflectivity is slightly higher over land, while the 99th percentile reflectivity is much higher 741 over land, indicating that land convection reaches more extreme values of intensity. The secondary 742 reflectivity maximum in the median line is also more noticeable over land than ocean, suggest-743 ing again significant differences in the vertical profile of the cloud systems over land and ocean. 744 Frequency is normalized at each level and all rain types (stratiform, convective, shallow isolated, 745 shallow non-isolated) are included. The bottom panels show how ascent profiles in rain systems 746 have much larger diurnal variations over western Africa and are much more top-heavy than over 747 the adjacent ocean. Profiles of vertical velocity in pressure coordinates (negative values indicate 748 upward motion) were obtained from the MERRA reanalysis. 749

The TRMM reflectivity data (Kummerow et al, 1998) was originally sourced from TRMM orbital files (2A23/2A25) for the 1998-2014 period (August values only). Vertical omega profiles are calculated from 6-hourly MERRA reanalyis (Rieckner et al, 2011) for August days in 1983-2007. Only samples that contributed the top 50% of rainfall by volume were included (using the MERRA surface precipitation flux data for all rain and rainfall thresholds unique to each domain and 6-hourly period). When all times are included, the eastern Atlantic ascent is even more bottom heavy than shown in Figure 4, but the qualitative comparison to land convection is unchanged.

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